

Molecular properties

Trond Saue

Laboratoire de Chimie et Physique Quantiques
CNRS/Université de Toulouse (Paul Sabatier)
118 route de Narbonne, 31062 Toulouse (FRANCE)
e-mail: trond.sau@irsamc.ups-tlse.fr



Maxwell's equations

for electric field \mathbf{E} and magnetic field \mathbf{B} in terms of sources ρ and \mathbf{j}

- The homogeneous pair:

$$\begin{aligned}\nabla \cdot \mathbf{B} &= 0 \\ \nabla \times \mathbf{E} + \partial_t \mathbf{B} &= 0; \quad \partial_t \mathbf{B} = \frac{\partial \mathbf{B}}{\partial t}\end{aligned}$$

- The inhomogeneous pair (sources):

$$\begin{aligned}\nabla \cdot \mathbf{E} &= \rho/\epsilon_0 \\ \nabla \times \mathbf{B} - \frac{1}{c^2} \partial_t \mathbf{E} &= \mu_0 \mathbf{j}\end{aligned}$$

- Electric and magnetic constants:

$$\mu_0 \epsilon_0 = \frac{1}{c^2}$$

- Introducing electromagnetic potentials solves the homogeneous pair

$$\mathbf{E} = -\nabla\phi - \frac{\partial \mathbf{A}}{\partial t}; \quad \mathbf{B} = \nabla \times \mathbf{A}$$

Maxwell's equations: The stationary case

- The homogeneous pair:

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = 0$$

- The inhomogeneous pair:

$$\nabla \cdot \mathbf{E} = \rho/\epsilon_0$$

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{j}$$

- Implies steady currents:

$$\nabla \cdot \mathbf{j} = 0 = -\partial_t \rho$$

- A useful formula:

$$\nabla \times (\nabla \times \mathbf{F}) = \nabla (\nabla \cdot \mathbf{F}) - \nabla^2 \mathbf{F}$$

Maxwell's equations: The stationary case

Electrostatics

$$\nabla^2 \mathbf{E} = \nabla \underbrace{(\nabla \cdot \mathbf{E})}_{\rho/\epsilon_0} - \nabla \times \underbrace{(\nabla \times \mathbf{E})}_{=0} = \nabla \rho / \epsilon_0$$



Siméon Denis Poisson
(1781-1840)

Each component of the electric field fulfills the Poisson equation:

$$\nabla^2 \Psi(\mathbf{r}_1, t) = f(\mathbf{r}_1, t)$$

with solutions

$$\Psi(\mathbf{r}_1, t) = -\frac{1}{4\pi} \int \frac{f(\mathbf{r}_2, t)}{r_{12}} d^3\mathbf{r}_2$$

Maxwell's equations: The stationary case

Electrostatics

$$\mathbf{E}(\mathbf{r}_1) = -\frac{1}{4\pi\epsilon_0} \int \frac{\nabla_2 \rho(\mathbf{r}_2)}{r_{12}} d^3 \mathbf{r}_2 = \frac{1}{4\pi\epsilon_0} \int \frac{\mathbf{r}_{12} \rho(\mathbf{r}_2)}{r_{12}^3} d^3 \mathbf{r}_2$$



Charles-Augustin de Coulomb
(1736 - 1806)

Coulomb's law:

$$\begin{aligned} \mathbf{F}(\mathbf{r}_1) &= q_1 \mathbf{E}(\mathbf{r}_1) \\ &= \frac{q_1}{4\pi\epsilon_0} \int \frac{(\mathbf{r}_1 - \mathbf{r}') \rho(\mathbf{r}')}{|\mathbf{r}_1 - \mathbf{r}'|^3} d^3 \mathbf{r}' \end{aligned}$$

Point charge:

$$\rho(\mathbf{r}') = q_2 \delta(\mathbf{r}' - \mathbf{r}_2) \Rightarrow \mathbf{F}(\mathbf{r}_1) = \frac{q_1 q_2 \mathbf{r}_{12}}{4\pi\epsilon_0 r_{12}^3}$$

Scalar form:

$$\mathbf{r}_{12} = r_{12} \mathbf{n}_{12} \rightarrow F = \frac{q_1 q_2}{4\pi\epsilon_0 r_{12}^2}$$

Maxwell's equations: The stationary case

Magnetostatics

$$\nabla^2 \mathbf{B} = \underbrace{\nabla(\nabla \cdot \mathbf{B})}_{=0} - \underbrace{\nabla \times (\nabla \times \mathbf{B})}_{\mu_0 \mathbf{j}} = -\mu_0 (\nabla \times \mathbf{j})$$



Jean-Baptiste Biot

(1774-1862)

Biot-Savart law:

$$\mathbf{B}(\mathbf{r}_1) = -\frac{\mu_0}{4\pi} \int \frac{\mathbf{r}_{12} \times \mathbf{j}(\mathbf{r}_2)}{r_{12}^3} d^3 \mathbf{r}_2$$

It would be tempting to insert the expression for a moving point charge

$$\mathbf{j}(\mathbf{r}') = q_2 v_2 \delta(\mathbf{r}' - \mathbf{r}_2(t))$$

but this is **wrong**, since a moving charge is not a steady current.

- From classical mechanics

$$H(\mathbf{r}, \mathbf{p}, t) = H_p(\mathbf{r}, \boldsymbol{\pi}, t) + q\phi(\mathbf{r}, t); \quad \mathbf{p} = \boldsymbol{\pi} + q\mathbf{A}$$

- Minimal substitution ($q = -e$):

$$\hat{H} \rightarrow \hat{H} - e\phi; \quad \hat{\mathbf{p}} \rightarrow \hat{\boldsymbol{\pi}} = \hat{\mathbf{p}} + e\mathbf{A}$$

- This **relativistic** coupling of particles and fields is also used in the non-relativistic domain.

Non-relativistic Hamiltonian in external fields

- Minimal substitution gives

$$h_0^{NR} = \frac{\hat{p}^2}{2m} \quad \rightarrow \quad h^{NR} = \frac{\hat{\pi}^2}{2m} - e\phi = \frac{\hat{p}^2}{2m} + \frac{e}{2m} [\hat{p} \cdot \mathbf{A} + \mathbf{A} \cdot \hat{p}] + \frac{e^2 A^2}{2m} - e\phi$$

- ▶ no spin interactions

- The Dirac identity

$$(\boldsymbol{\sigma} \cdot \mathbf{A})(\boldsymbol{\sigma} \cdot \mathbf{B}) = \mathbf{A} \cdot \mathbf{B} + i\boldsymbol{\sigma} \cdot (\mathbf{A} \times \mathbf{B})$$

- A special case

$$(\boldsymbol{\sigma} \cdot \hat{\mathbf{p}})(\boldsymbol{\sigma} \cdot \hat{\mathbf{p}}) = \hat{p}^2$$

suggests that spin is “hidden” in the non-relativistic operator.

- Minimal substitution then gives

$$\begin{aligned} h_0^{NR} = \frac{(\boldsymbol{\sigma} \cdot \hat{\mathbf{p}})^2}{2m} \quad \rightarrow \quad h^{NR} &= \frac{(\boldsymbol{\sigma} \cdot \hat{\boldsymbol{\pi}})^2}{2m} - e\phi \\ &= \frac{\hat{p}^2}{2m} + \frac{e}{2m} [\hat{p} \cdot \mathbf{A} + \mathbf{A} \cdot \hat{p}] + \frac{e^2 A^2}{2m} + \frac{e\hbar}{2m} (\boldsymbol{\sigma} \cdot \mathbf{B}) - e\phi \end{aligned}$$

- Minimal substitution gives

$$h_0^R = \beta mc^2 + c(\boldsymbol{\alpha} \cdot \mathbf{p}) \quad \rightarrow \quad h^R = \beta mc^2 + c(\boldsymbol{\alpha} \cdot \mathbf{p}) + ec(\boldsymbol{\alpha} \cdot \mathbf{A}) - e\phi$$

- The expectation value of the interaction Hamiltonian is given in terms of the electromagnetic potentials

$$\langle H_{int} \rangle = \int [\rho(\mathbf{r})\phi(\mathbf{r}) - \mathbf{j}(\mathbf{r}) \cdot \mathbf{A}(\mathbf{r})] d^3\mathbf{r}$$

- Is it possible to express the interaction Hamiltonian directly in terms of electromagnetic fields ?
 - ▶ The answer is: **Yes**,
 - ▶ using **multipolar gauge**

Multipolar gauge

Taylor expansion of electromagnetic potentials

- Scalar potential:

$$\begin{aligned}\tilde{\phi}(\mathbf{r}, t) &= \tilde{\phi}(\mathbf{a}, t) + \left[(\boldsymbol{\delta} \cdot \nabla') \tilde{\phi}(\mathbf{r}', t) \right]_{\mathbf{r}'=\mathbf{a}} + \dots \\ &= \sum_{n=0}^{\infty} \frac{1}{n!} \left[(\boldsymbol{\delta} \cdot \nabla')^n \tilde{\phi}(\mathbf{r}', t) \right]_{\mathbf{r}'=\mathbf{a}}; \quad \boldsymbol{\delta} = \mathbf{r} - \mathbf{a}\end{aligned}$$

where \mathbf{a} is the expansion point.

- ▶ We shall set $\mathbf{a} = \mathbf{0}$, such that $\boldsymbol{\delta} = \mathbf{r}$.

- Likewise, for the vector potential

$$\tilde{\mathbf{A}}(\mathbf{r}, t) = \sum_{n=0}^{\infty} \frac{1}{n!} \left[(\mathbf{r} \cdot \nabla')^n \tilde{\mathbf{A}}(\mathbf{r}', t) \right]_{\mathbf{r}'=\mathbf{0}}$$

Multipolar gauge

Replacing potentials by fields

- Using the relation $\mathbf{E} = -\nabla\phi - \frac{\partial\mathbf{A}}{\partial t}$ we obtain

$$\begin{aligned}\tilde{\phi}(\mathbf{r}, t) &= \tilde{\phi}(\mathbf{0}, t) - \sum_{n=1}^{\infty} \frac{1}{n!} \left[(\mathbf{r} \cdot \nabla')^{n-1} (\mathbf{r} \cdot \mathbf{E}(\mathbf{r}', t)) \right]_{\mathbf{r}'=\mathbf{0}} \\ &\quad - \frac{\partial}{\partial t} \sum_{n=1}^{\infty} \frac{1}{n!} \left[(\mathbf{r} \cdot \nabla')^{n-1} (\mathbf{r} \cdot \tilde{\mathbf{A}}(\mathbf{r}', t)) \right]_{\mathbf{r}'=\mathbf{0}}\end{aligned}$$

- which can be written as a gauge transformation

$$\tilde{\phi}(\mathbf{r}, t) = \phi(\mathbf{r}, t) - \frac{\partial}{\partial t} \chi(\mathbf{r}, t)$$

- with the gauge function

$$\chi(\mathbf{r}, t) = \sum_{n=1}^{\infty} \frac{1}{n!} \left[(\mathbf{r} \cdot \nabla')^{n-1} (\mathbf{r} \cdot \tilde{\mathbf{A}}(\mathbf{r}', t)) \right]_{\mathbf{r}'=\mathbf{0}}$$

- We next carry out the gauge transformation

$$\tilde{\mathbf{A}}(\mathbf{r}, t) = \mathbf{A}(\mathbf{r}, t) + \nabla\chi(\mathbf{r}, t)$$

- After some further manipulations we arrive at the final expressions

$$\phi(\mathbf{r}, t) = \tilde{\phi}(\mathbf{0}, t) - \sum_{n=0}^{\infty} \frac{1}{(n+1)!} [(\mathbf{r} \cdot \nabla')^n (\mathbf{r} \cdot \mathbf{E}(\mathbf{r}', t))]_{\mathbf{r}'=\mathbf{0}}$$

$$\mathbf{A}(\mathbf{r}, t) = - \sum_{n=1}^{\infty} \frac{n}{(n+1)!} [(\mathbf{r} \cdot \nabla')^{n-1} (\mathbf{r} \times \mathbf{B}(\mathbf{r}', t))]_{\mathbf{r}'=\mathbf{0}}$$

- What happened to gauge freedom ?

Multipolar gauge

Examples

$$\phi(\mathbf{r}, t) = \phi^{[0]} - \sum_{n=0}^{\infty} \frac{1}{(n+1)!} r_{j_1} r_{j_2} \cdots r_{j_n} \left(\mathbf{r} \cdot \mathbf{E}_{j_1 j_2 \dots j_n}^{[n]} \right)$$

$$\mathbf{A}(\mathbf{r}, t) = - \sum_{n=0}^{\infty} \frac{n+1}{(n+2)!} r_{j_1} r_{j_2} \cdots r_{j_n} \left(\mathbf{r} \times \mathbf{B}_{j_1 j_2 \dots j_n}^{[n]} \right)$$

- Uniform electric field:

$$\phi(\mathbf{r}, t) = -\mathbf{r} \cdot \mathbf{E}^{[0]}; \quad \mathbf{A}(\mathbf{r}, t) = \mathbf{0}$$

- Uniform magnetic field:

$$\phi(\mathbf{r}, t) = 0; \quad \mathbf{A}(\mathbf{r}, t) = \frac{1}{2} \left(\mathbf{B}^{[0]} \times \mathbf{r} \right)$$

- For a **time-dependent** uniform magnetic field we get a non-uniform electric field as well

$$\mathbf{E}(\mathbf{r}, t) = -\nabla\phi - \partial_t \mathbf{A} = -\frac{1}{2} \left(\mathbf{r} \times \partial_t \mathbf{B}^{[0]} \right)$$

Multipole expansions

- In multipolar gauge the expectation value of the interaction Hamiltonian takes the form

$$\begin{aligned}\langle H_{int} \rangle &= \int [\rho(\mathbf{r})\phi(\mathbf{r}) - \mathbf{j}(\mathbf{r}) \cdot \mathbf{A}(\mathbf{r})] d^3\mathbf{r} \\ &= Q^{[0]}\phi^{[0]} - \sum_{n=1}^{\infty} \frac{1}{n!} \mathbf{Q}_{j_1 j_2 \dots j_{n-1}}^{[n]} \cdot \mathbf{E}_{j_1 \dots j_{n-1}}^{[n-1]} - \sum_{n=1}^{\infty} \frac{1}{n!} \mathbf{m}_{j_1 j_2 \dots j_{n-1}}^{[n]} \cdot \mathbf{B}_{j_1 \dots j_{n-1}}^{[n-1]}\end{aligned}$$

- where appears electric multipoles

$$Q_{j_1 \dots j_n}^{[n]} = \int r_{j_1} r_{j_2} \dots r_{j_n} \rho(\mathbf{r}) d^3\mathbf{r}$$

- and magnetic multipoles

$$\mathbf{m}_{j_1 \dots j_{n-1}}^{[n]} = \frac{n}{n+1} \int r_{j_1} r_{j_2} \dots r_{j_{n-1}} [\mathbf{r} \times \mathbf{j}(\mathbf{r})] d^3\mathbf{r}$$

External fields

in terms of electromagnetic potentials

- Uniform electric field

$$\phi(\mathbf{r}_i) = -\mathbf{r}_i \cdot \mathbf{E}$$

- Uniform magnetic field

$$\mathbf{A}(\mathbf{r}_i) = \frac{1}{2} (\mathbf{B} \times \mathbf{r}_{iG})$$

- Nuclear spin:

$$\mathbf{m}_K = \gamma \mathbf{l}_K; \quad \gamma - \text{gyromagnetic ratio}$$

- Vector potential of point-like nuclear magnetic dipole

$$\mathbf{A}_K(\mathbf{r}_i) = \frac{\mu_0}{4\pi} \frac{\mathbf{m}_K \times \mathbf{r}_{iK}}{r_{iK}^3}$$

- Corresponding magnetic field

$$\mathbf{B}_K(\mathbf{r}_i) = \nabla \times \mathbf{A}_K = \frac{\mu_0}{4\pi} \left[\mathbf{m}_K \frac{8\pi}{3} \delta(\mathbf{r}_{iK}) - \frac{\mathbf{m}_K r_{iK}^2 - 3\mathbf{r}_{iK}(\mathbf{r}_{iK} \cdot \mathbf{m}_K)}{r_{iK}^3} \right]$$

- General form:

$$\hat{h}_{int}^R = ec(\boldsymbol{\alpha} \cdot \mathbf{A}) - e\phi$$

- Uniform electric field:

$$\hat{h}_{E1} = -\boldsymbol{\mu} \cdot \mathbf{E}; \quad \boldsymbol{\mu} = -e\mathbf{r}$$

- Uniform magnetic field (Zeeman interaction):

$$\hat{h}_{rel}^Z = -\mathbf{m}_e \cdot \mathbf{B}; \quad \mathbf{m}_e(\mathbf{r}_i) = \frac{1}{2}(\mathbf{r}_{iG} \times \hat{\mathbf{j}}_{rel}); \quad \hat{\mathbf{j}}_{rel} = -ec\boldsymbol{\alpha}$$

- Nuclear spins:

$$\hat{h}_{rel}^{hfs} = -\sum_K \mathbf{m}_K \cdot \hat{\mathbf{B}}_K^{el}; \quad \hat{\mathbf{B}}_K^{el}(\mathbf{r}_i) = -\frac{\mu_0}{4\pi} \frac{\mathbf{r}_{iK} \times ec\boldsymbol{\alpha}}{r_{iK}^3}$$

Non-relativistic interaction Hamiltonian

- General form:

$$\hat{h}_{int}^{NR} = \frac{e}{2m} [\hat{\mathbf{p}} \cdot \mathbf{A} + \mathbf{A} \cdot \hat{\mathbf{p}}] + \frac{e^2 A^2}{2m} + \frac{e\hbar}{2m} (\boldsymbol{\sigma} \cdot \mathbf{B}) - e\phi$$

- Coulomb gauge ($\nabla \cdot \mathbf{A} = 0$)

$$\hat{h}_{int}^{NR} = \frac{e}{m} (\mathbf{A} \cdot \hat{\mathbf{p}}) + \frac{e^2 A^2}{2m} + \frac{e}{2m} (\boldsymbol{\sigma} \cdot \mathbf{B}) - e\phi$$

- Spin interaction:

$$\hat{h} = \frac{e\hbar}{2m} (\boldsymbol{\sigma} \cdot \mathbf{B}) \quad \rightarrow \quad \frac{e}{m} (\hat{\mathbf{s}} \cdot \mathbf{B}) \quad \rightarrow \quad \frac{g_e e}{2m} (\hat{\mathbf{s}} \cdot \mathbf{B})$$

- ▶ Electronic g-factor

$$g_e = 2.0023193043617(15)$$

- Uniform electric field:

$$\hat{h}_{E1} = -\boldsymbol{\mu} \cdot \mathbf{E}; \quad \boldsymbol{\mu} = -e\mathbf{r}$$

Non-relativistic interaction Hamiltonian

Zeeman interaction

- General form:

$$\hat{h}_{int}^{NR} = \frac{e}{m} (\mathbf{A} \cdot \hat{\mathbf{p}}) + \frac{e^2 A^2}{2m} + \frac{g_e e}{2m} (\hat{\mathbf{s}} \cdot \mathbf{B}) - e\phi$$

- Uniform magnetic field (Zeeman interaction):

- ▶ Orbital Zeeman:

$$\hat{h}^{OZ} = \frac{e}{2m} \hat{\ell}_G \cdot \mathbf{B}; \quad \hat{\ell}_G(i) = \mathbf{r}_{iG} \times \hat{\mathbf{p}}$$

- ▶ Spin Zeeman:

$$\hat{h}^{SZ} = \frac{g_e e}{2m} \hat{\mathbf{s}} \cdot \mathbf{B}$$

- ▶ Total paramagnetic contribution:

$$\hat{h}^Z = -\mathbf{m}^{para} \cdot \mathbf{B}; \quad \mathbf{m}^{para}(i) = -\frac{e}{2m} (\hat{\ell}_G(i) + g_e \hat{\mathbf{s}}(i))$$

- ▶ Diamagnetic contribution:

$$\hat{h}_{BB}^{dia} = \frac{e^2}{8m} [B^2 r_{iG}^2 - (\mathbf{B} \cdot \mathbf{r}_{iG})(\mathbf{r}_{iG} \cdot \mathbf{B})]$$

Non-relativistic interaction Hamiltonian

Hyperfine interaction

- General form:

$$\hat{h}_{int}^{NR} = \frac{e}{m} (\mathbf{A} \cdot \hat{\mathbf{p}}) + \frac{e^2 A^2}{2m} + \frac{g_e e}{2m} (\hat{\mathbf{s}} \cdot \mathbf{B}) - e\phi$$

- Nuclear spins:

- ▶ Orbital contributions: $\sum_K \hat{h}_K^{psO} + \sum_{KL} \hat{h}^{dso}$

- ★ Paramagnetic spin-orbit:

$$\hat{h}_K^{psO} = \frac{\mu_0}{4\pi} \frac{e}{m r_{iK}^3} \hat{\mathbf{m}}_K \cdot \hat{\mathbf{l}}_K$$

- ★ Diamagnetic spin-orbit:

$$\hat{h}_{KL}^{dso} = \left(\frac{e}{2m}\right)^2 \left(\frac{\mu_0}{4\pi}\right)^2 \left[\frac{(\mathbf{m}_K \cdot \mathbf{m}_L)(\mathbf{r}_{iK} \cdot \mathbf{r}_{iL}) - (\mathbf{m}_K \cdot \mathbf{r}_{iL})(\mathbf{r}_{iK} \cdot \mathbf{m}_L)}{r_{iK}^3 r_{iL}^3} \right]$$

- ▶ Spin contributions: $\sum_K (\hat{h}_K^{fc} + \hat{h}_K^{sd})$

- ★ Fermi contact (near-field):

$$\hat{h}_K^{fc} = \left(\frac{\mu_0}{4\pi}\right) \left(\frac{g_e e}{2m}\right) \left(\hat{\mathbf{s}} \cdot \mathbf{m}_K \frac{8\pi}{3} \delta(\mathbf{r}_{iK})\right)$$

- ★ Spin-dipolar term (far-field):

$$\hat{h}_K^{sd} = - \left(\frac{\mu_0}{4\pi}\right) \left(\frac{g_e e}{2m}\right) \hat{\mathbf{s}} \cdot \left[\frac{\mathbf{m}_K r_{iK}^2 - 3r_{iK} (\mathbf{r}_{iK} \cdot \mathbf{m}_K)}{r_{iK}^3} \right]$$

Non-relativistic interaction operator

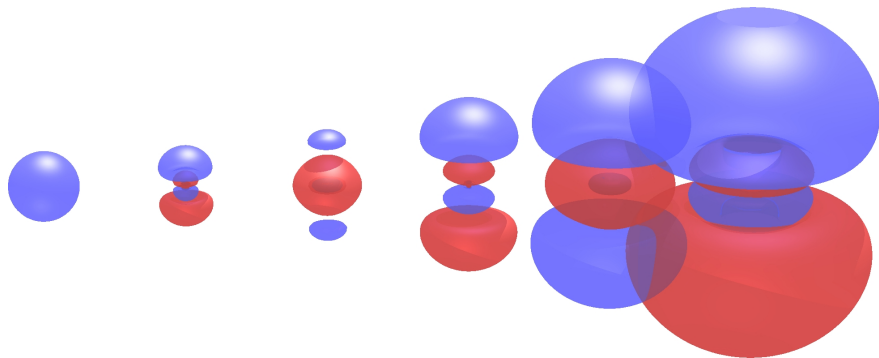
The NMR experiment

- The NMR experiment involves an external magnetic field as well as nuclear spins

$$\hat{h}_{int}^{NR} = \hat{h}^{OZ} + \hat{h}^{SZ} + \hat{h}_{BB}^{dia} \sum_K \left(\hat{h}_K^{psO} + \hat{h}_K^{fc} + \hat{h}_K^{sd} + \hat{h}_{BK}^{dia} \right) + \sum_{KL} \hat{h}^{dso},$$

- where appears the mixed diamagnetic contribution

$$\hat{h}_{BK}^{dia} = \frac{\mu_0}{4\pi} \frac{e}{2m} \left[\left[\frac{(\mathbf{B} \cdot \mathbf{m}_K)(\mathbf{r}_{iG} \cdot \mathbf{r}_{iK}) - (\mathbf{B} \cdot \mathbf{r}_{iK})(\mathbf{r}_{iG} \cdot \mathbf{m}_K)}{r_{iK}^3} \right] \right]$$



- $\rho = \rho^{(0)} + \rho^{(1)}F_z + \frac{1}{2!}\rho^{(2)}F_z^2 + \frac{1}{3!}\rho^{(3)}F_z^3 + \frac{1}{4!}\rho^{(4)}F_z^4 + \frac{1}{5!}\rho^{(5)}F_z^5 + \dots$
- $\mu_i = \int r_i \rho^{(0)} d\tau + \underbrace{\int r_i \rho^{(z)} d\tau}_{\alpha_{iz}} F_z + \frac{1}{2!} \underbrace{\int r_i \rho^{(zz)} d\tau}_{\beta_{izz}} F_z F_z + \dots$

Response functions

- Consider a Hamiltonian on the form:

$$\hat{H} = \hat{H}_0 + \hat{V}(t); \quad \hat{V}(t) = \int_{-\infty}^{+\infty} \hat{V}(\omega) e^{-i\omega t} d\omega$$

- Hermiticity implies:

$$\hat{V}^\dagger(t) = \hat{V}(t) \quad \Rightarrow \quad \hat{V}^\dagger(\omega) = \hat{V}(-\omega)$$

- Kubo expansion:

$$\begin{aligned} \langle \tilde{0} | \hat{\Omega} | \tilde{0} \rangle &= \langle 0 | \hat{\Omega} | 0 \rangle \\ &+ \int_{-\infty}^{+\infty} \langle \langle \hat{\Omega}; \hat{V}(\omega) \rangle \rangle e^{-i\omega t} d\omega \\ &+ \frac{1}{2} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \langle \langle \hat{\Omega}; \hat{V}(\omega_1), \hat{V}(\omega_2) \rangle \rangle e^{-i(\omega_1 + \omega_2)t} d\omega_1 d\omega_2 \\ &+ \frac{1}{6} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \langle \langle \hat{\Omega}; \hat{V}(\omega_1), \hat{V}(\omega_2), \hat{V}(\omega_3) \rangle \rangle e^{-i(\omega_1 + \omega_2 + \omega_3)t} d\omega_1 d\omega_2 d\omega_3 \\ &+ \dots \end{aligned}$$

- Exact states: the connection between expectation values and energy derivatives is provided by the Hellmann-Feynmann theorem.

- We start from an energy function

$$E \equiv E(\lambda, \epsilon) \quad \begin{cases} \lambda & \text{(variational parameters)} \\ \epsilon & \text{(perturbation strengths)} \end{cases}$$

- We have to carefully distinguish total and partial derivatives.
Consider the differential

$$dE = \sum_X \left(\frac{\partial E}{\partial \epsilon_X} \right) d\epsilon_X + \sum_i \left(\frac{\partial E}{\partial \lambda_i} \right) d\lambda_i$$

- Total derivative

$$\begin{aligned} \frac{dE}{d\epsilon_A} &= \sum_X \left(\frac{\partial E}{\partial \epsilon_X} \right) \frac{d\epsilon_X}{d\epsilon_A} + \sum_i \left(\frac{\partial E}{\partial \lambda_i} \right) \frac{d\lambda_i}{d\epsilon_A} \\ &= \frac{\partial E}{\partial \epsilon_A} + \sum_i \left(\frac{\partial E}{\partial \lambda_i} \right) \frac{d\lambda_i}{d\epsilon_A} \end{aligned}$$

Variational perturbation theory

- We start from

$$\left[\frac{dE}{d\varepsilon_A} \right]_{\varepsilon=0} = \left[\frac{\partial E}{\partial \varepsilon_A} + \sum_i \left(\frac{\partial E}{\partial \lambda_i} \right) \frac{d\lambda_i}{d\varepsilon_A} \right]_{\varepsilon=0}$$

- Key assumption of variational perturbation theory

$$\left. \frac{\partial E}{\partial \lambda_i} \right|_{\varepsilon} = 0, \quad \forall \lambda_i$$

- We then get

$$\left[\frac{dE}{d\varepsilon_A} \right]_{\varepsilon=0} = \left[\frac{\partial E}{\partial \varepsilon_A} \right]_{\varepsilon=0}$$

- Second derivatives

$$\left[\frac{d^2 E}{d\varepsilon_A d\varepsilon_B} \right]_{\varepsilon=0} = \left[\frac{\partial^2 E}{\partial \varepsilon_A \partial \varepsilon_B} + \sum_i \frac{\partial^2 E}{\partial \varepsilon_A \partial \lambda_i} \frac{\partial \lambda_i}{\partial \varepsilon_B} \right]_{\varepsilon=0},$$

- ..translated to vector notation

$$\left[\frac{d^2 E}{d\varepsilon_A d\varepsilon_B} \right]_{\varepsilon=0} = E_{AB}^{[0]} + \mathbf{E}_A^{[1]} \cdot \boldsymbol{\lambda}_B^{[1]}$$

- The variational condition implies

$$\left. \frac{\partial E}{\partial \lambda_i} \right|_{\epsilon} = 0, \quad \forall \lambda_i \quad \Rightarrow \quad \left. \frac{d}{d\epsilon_A} \frac{\partial E}{\partial \lambda_i} \right|_{\epsilon} = 0$$

- This allows us to generate response equations, e.g

$$\left[\frac{d}{d\epsilon_A} \left(\frac{\partial E}{\partial \lambda_i} \right) \right]_{\epsilon=0} = \left[\frac{\partial^2 E}{\partial \epsilon_A \partial \lambda_i} + \frac{\partial^2 E}{\partial \lambda_i \partial \lambda_j} \frac{\partial \lambda_j}{\partial \epsilon_A} \right]_{\epsilon=0} = 0$$

- ..translated into vector notation

$$\mathbf{E}_A^{[1]} + E_0^{[2]} \boldsymbol{\lambda}_A^{[1]} = 0; \quad \Rightarrow \quad \boldsymbol{\lambda}_A^{[1]} = - \left(E_0^{[2]} \right)^{-1} \mathbf{E}_A^{[1]}$$

- Second derivative

$$\left[\frac{d^2 E}{d\epsilon_A d\epsilon_B} \right]_{\epsilon=0} = E_{AB}^{[0]} + \mathbf{E}_A^{[1]} \cdot \boldsymbol{\lambda}_B^{[1]} = E_{AB}^{[0]} - \mathbf{E}_A^{[1]} \cdot \left(E_0^{[2]} \right)^{-1} \mathbf{E}_B^{[1]}$$

Exact-state response functions

- Let us consider the exact-state response function. It can be derived as a FCI problem.
- We consider a time-independent Hamiltonian on the form

$$\hat{H} = \hat{H}_0 + \hat{V}(\epsilon); \quad \hat{V}(\epsilon) = \sum_X \epsilon_X \hat{h}_X + \frac{1}{2} \sum_{X,Y} \epsilon_X \epsilon_Y \hat{h}_{XY}$$

- We assume that we have the exact solutions $\{|n\rangle\}$ of \hat{H}_0
- We write the perturbed wave-function as

$$|\mathbf{c}\rangle = \sum_n |n\rangle c_n \xrightarrow{\epsilon=0} |0\rangle$$

- We shall impose the normalization condition

$$\langle \mathbf{c} | \mathbf{c} \rangle = \sum_m \sum_n c_m \langle m | n \rangle c_n = \sum_n c_n^2 = 1$$

- The energy is

$$E^{FCI}(\mathbf{c}, \epsilon) = \langle \mathbf{c} | \hat{H} | \mathbf{c} \rangle = \sum_m \sum_n c_m \langle m | \hat{H} | n \rangle c_n \xrightarrow{\epsilon=0} \langle 0 | \hat{H} | 0 \rangle = E_0$$

Exact-state response functions

- We set up a Lagrangian

$$L^{FCI}(\mathbf{c}, \varepsilon, \mu) = E^{FCI}(\mathbf{c}, \varepsilon) - \mu(\langle \mathbf{c} | \mathbf{c} \rangle - 1)$$

- We impose variational conditions

- ▶ on the CI-coefficients

$$\left. \frac{\partial L^{FCI}}{\partial c_n} \right|_{\varepsilon} = \frac{\partial E^{FCI}(\mathbf{c}, \varepsilon)}{\partial c_n} - \mu \frac{\partial \langle \mathbf{c} | \mathbf{c} \rangle}{\partial c_n} = 2 \langle n | \hat{H} - \mu | \mathbf{c} \rangle = 0$$

- ★ giving us a FCI-problem

$$H\mathbf{c} = \mu\mathbf{c}; \quad H_{mn} = \langle m | \hat{H} | n \rangle$$

- ▶ on the Lagrange-multiplier

$$\left. \frac{\partial L^{FCI}}{\partial \mu} \right|_{\varepsilon} = 1 - \langle \mathbf{c} | \mathbf{c} \rangle = 0$$

- ★ ..returning the normalization condition.

First-order response equations

- Since the variational conditions are to hold at *any* perturbation strength, their total derivatives with respect to perturbation strength are zero at *any* perturbation strength.
- We shall need the first-order response equation from the CI-coefficients:

$$\begin{aligned}
 0 &= \left[\frac{d}{d\varepsilon_B} \left(\frac{\partial L^{FCI}}{\partial c_n} \right) \right]_{\varepsilon=0} = \left[\frac{\partial^2 L^{FCI}}{\partial c_n \partial \varepsilon_B} + \sum_m \frac{\partial^2 L^{FCI}}{\partial c_n \partial c_m} \frac{dc_m}{d\varepsilon_B} + \frac{\partial^2 L^{FCI}}{\partial c_n \partial \mu} \frac{d\mu}{d\varepsilon_B} \right]_{\varepsilon=0} \\
 &= \left[\langle n | \frac{\partial \hat{H}}{\partial \varepsilon_B} | \mathbf{c} \rangle + \sum_m 2 \langle n | \hat{H} - \mu | m \rangle \frac{dc_m}{d\varepsilon_B} + 2 \langle n | \mathbf{c} \rangle \frac{d\mu}{d\varepsilon_B} \right]_{\varepsilon=0} \\
 &= \langle n | \hat{h}_B | 0 \rangle + \sum_m 2 \langle n | \hat{H}_0 - E_0 | m \rangle c_m^{[B]} + 2 \langle n | 0 \rangle \mu^{[B]} \\
 &= \langle n | \hat{h}_B | 0 \rangle + 2 (E_n - E_0) c_n^{[B]} + 2 \delta_{n0} \mu^{[B]} = 0
 \end{aligned}$$

- .. as well as the Lagrange multiplier

$$\begin{aligned}
 \left[\frac{d}{d\varepsilon_B} \left(\frac{\partial L^{FCI}}{\partial \mu} \right) \right]_{\varepsilon=0} &= \left[\underbrace{\frac{\partial^2 L^{FCI}}{\partial \mu \partial \varepsilon_B}}_{=0} + \sum_m \frac{\partial^2 L^{FCI}}{\partial \mu \partial c_m} \frac{dc_m}{d\varepsilon_B} + \underbrace{\frac{\partial^2 L^{FCI}}{\partial \mu \partial \mu}}_{=0} \frac{d\mu}{d\varepsilon_B} \right]_{\varepsilon=0} \\
 &= \left[- \sum_m 2 \langle \mathbf{c} | m \rangle \frac{dc_m}{d\varepsilon_B} \right]_{\varepsilon=0} = - \sum_m 2 \langle 0 | m \rangle c_m^{[B]} = -2c_0^{[B]} = 0
 \end{aligned}$$

- We first consider

$$\left. \frac{dL^{FCI}}{d\varepsilon_A} \right|_{\varepsilon} = \left[\frac{\partial L^{FCI}}{\partial \varepsilon_A} + \sum_n \frac{\partial L^{FCI}}{\partial c_n} \frac{dc_n}{d\varepsilon_A} + \frac{\partial L^{FCI}}{\partial \mu} \frac{d\mu}{d\varepsilon_A} \right]_{\varepsilon} = \left[\frac{\partial L^{FCI}}{\partial \varepsilon_A} \right]_{\varepsilon},$$

▶ where we used the variational conditions.

- Specifically, we calculate

$$\left. \frac{dL^{FCI}}{d\varepsilon_A} \right|_{\varepsilon=0} = \left[\frac{\partial L^{FCI}}{\partial \varepsilon_A} \right]_{\varepsilon=0} = \left[\langle \mathbf{c} | \frac{\partial \hat{H}}{\partial \varepsilon_A} | \mathbf{c} \rangle \right]_{\varepsilon=0} = \langle 0 | \hat{h}_A | 0 \rangle$$

- We continue to the second derivative

$$\begin{aligned}
 \left. \frac{dL^{FCI}}{d\varepsilon_A d\varepsilon_B} \right|_{\varepsilon=0} &= \left[\frac{d}{d\varepsilon_B} \left(\frac{\partial L^{FCI}}{\partial \varepsilon_A} \right) \right]_{\varepsilon=0} \\
 &= \left[\frac{\partial^2 L^{FCI}}{\partial \varepsilon_A \partial \varepsilon_B} + \sum_n \frac{\partial^2 L^{FCI}}{\partial \varepsilon_A \partial c_n} \frac{dc_n}{d\varepsilon_B} + \underbrace{\frac{\partial^2 L^{FCI}}{\partial \varepsilon_A \partial \mu}}_{=0} \frac{d\mu}{d\varepsilon_B} \right]_{\varepsilon=0} \\
 &= \left[\langle \mathbf{c} | \frac{\partial^2 \hat{H}}{\partial \varepsilon_A \partial \varepsilon_B} | \mathbf{c} \rangle + \sum_n 2 \langle \mathbf{c} | \frac{\partial \hat{H}}{\partial \varepsilon_A} | n \rangle \frac{dc_n}{d\varepsilon_B} \right]_{\varepsilon=0} \\
 &= \langle 0 | \hat{h}_{AB} | 0 \rangle + \sum_n 2 \langle 0 | \hat{h}_A | n \rangle c_n^{[B]}
 \end{aligned}$$

- From the response equations we know that $c_0^{[B]} = 0$ and

$$c_{n \neq 0}^{[B]} = -\frac{1}{2} \frac{\langle n | \hat{h}_B | 0 \rangle}{E_n - E_0} \Rightarrow \left. \frac{dL^{FCI}}{d\varepsilon_A d\varepsilon_B} \right|_{\varepsilon=0} = \langle 0 | \hat{h}_{AB} | 0 \rangle - \sum_{n \neq 0} \frac{\langle 0 | \hat{h}_A | n \rangle \langle n | \hat{h}_B | 0 \rangle}{E_n - E_0}$$

- We recover the expression from Rayleigh-Schrödinger perturbation theory
- For second-order NR magnetic properties the diamagnetic contribution is an expectation value, the paramagnetic one a linear response function.
- The linear response function contains excitation energies and transition moments.

Hartree-Fock response theory

BCH-expansion of HF-energy

- We write the perturbed HF state as

$$|\tilde{0}\rangle = \exp(-\hat{\kappa})|0\rangle; \quad \hat{\kappa} = \sum_{pq} \kappa_{pq} \hat{a}_p^\dagger \hat{a}_q,$$

where $|0\rangle$ is the unperturbed HF solution, obtained with \hat{H}_0 .

- No Lagrange-multipliers are needed !
- The perturbed HF energy is

$$E^{HF}(\kappa) = \langle \tilde{0} | \hat{H} | \tilde{0} \rangle = \langle 0 | \exp(\hat{\kappa}) \hat{H} \exp(-\hat{\kappa}) | 0 \rangle$$

- We obtain an expansion in order of κ using the BCH-expansion

$$\begin{aligned} E^{HF}(\kappa) &= \langle 0 | \hat{H} | 0 \rangle && O(\kappa^0) \\ &+ \langle 0 | [\hat{\kappa}, \hat{H}] | 0 \rangle && O(\kappa^1) \\ &+ \frac{1}{2} \langle 0 | [\hat{\kappa}, [\hat{\kappa}, \hat{H}]] | 0 \rangle && O(\kappa^2) \\ &+ \frac{1}{6} \langle 0 | [\hat{\kappa}, [\hat{\kappa}, [\hat{\kappa}, \hat{H}]]] | 0 \rangle && O(\kappa^3) \\ &+ \dots \end{aligned}$$

Hartree-Fock response theory

A very useful relation !

- Starting from

$$\hat{H} = \sum_{pq} h_{pq} \hat{a}_p^\dagger \hat{a}_q + \frac{1}{2} \sum_{pqrs} (pq|rs) \hat{a}_p^\dagger \hat{a}_r^\dagger \hat{a}_s \hat{a}_q$$

- .. one can derive

$$[\hat{\kappa}, \hat{H}] = \hat{H}^{\{1\}} = \sum_{pq} h_{pq}^{\{1\}} \hat{a}_p^\dagger \hat{a}_q + \frac{1}{2} \sum_{pqrs} (pq|rs)^{\{1\}} \hat{a}_p^\dagger \hat{a}_r^\dagger \hat{a}_s \hat{a}_q,$$

- ..where

$$h_{pq}^{\{1\}} = \sum_t (\kappa_{pt} h_{tq} - h_{pt} \kappa_{tq})$$
$$(pq|rs)^{\{1\}} = \sum_t [(\kappa_{pt} (tq|rs) - (pt|rs) \kappa_{tq}) + (\kappa_{rt} (tq|rs) - (pt|rs) \kappa_{ts})]$$

- ..such that

$$E^{HF}(\kappa) = \langle 0 | \hat{H} | 0 \rangle + \langle 0 | \hat{H}^{\{1\}} | 0 \rangle + \frac{1}{2} \langle 0 | [\hat{\kappa}, \hat{H}^{\{1\}}] | 0 \rangle + \frac{1}{6} \langle 0 | \left[\hat{\kappa}, \frac{1}{2} \langle 0 | [\hat{\kappa}, \hat{H}^{\{1\}}] | 0 \rangle \right] | 0 \rangle$$
$$+ \langle 0 | \hat{H} | 0 \rangle + \langle 0 | \hat{H}^{\{1\}} | 0 \rangle + \frac{1}{2} \langle 0 | \hat{H}^{\{2\}} | 0 \rangle + \frac{1}{6} \langle 0 | \hat{H}^{\{3\}} | 0 \rangle + \dots$$

Hartree-Fock response theory

Redundant variables

- All terms of the BCH-expanded HF-energy have the same form

$$\langle 0 | \hat{H}^{\{n\}} | 0 \rangle = \sum_i h_{ii}^{\{n\}} + \frac{1}{2} \sum_{ij} (ii||jj)^{\{n\}}; \quad (pq||rs) = \{pq|rs\} - (ps|rq)$$

- We shall remove redundant variational parameters; they do not contribute to $\langle 0 | \hat{H}^{\{1\}} | 0 \rangle$. We note

$$\begin{aligned} \sum_i h_{ii}^{\{1\}} &= \sum_i \sum_t (\kappa_{it} h_{ti} - h_{it} \kappa_{ti}) \\ &= \sum_i \sum_j (\kappa_{ij} h_{ji} - h_{ij} \kappa_{ji}) + \sum_i \sum_a (\kappa_{ia} h_{ai} - h_{ia} \kappa_{ai}) \end{aligned}$$

- The one-electron term involving only occupied orbitals is zero

$$\sum_i \sum_j (\kappa_{ij} h_{ji} - h_{ij} \kappa_{ji}) = \sum_i \sum_j \kappa_{ij} h_{ji} - \sum_i \sum_j h_{ij} \kappa_{ji} = \sum_i \sum_j \kappa_{ij} h_{ji} - \sum_i \sum_j h_{ji} \kappa_{ij} = 0$$

- The same holds for the two-electron term; orbital rotation parameters $\{\kappa_{ij}\}$ are redundant.

- The same holds for $\{\kappa_{ab}\}$ since they do not appear in these expressions at all.
- We can therefore write

$$\hat{\kappa} = \sum_{pq} \kappa_{pq} \hat{a}_p^\dagger \hat{a}_q \rightarrow \hat{\kappa} = \sum_{ai} \left[\kappa_{ai} \hat{a}_a^\dagger \hat{a}_i + \kappa_{ia} \hat{a}_i^\dagger \hat{a}_a \right] = \sum_{ai} \left[\kappa_{ai} \hat{a}_a^\dagger \hat{a}_i - \kappa_{ai}^* \hat{a}_i^\dagger \hat{a}_a \right]$$

Hartree-Fock response theory

HF gradient and first-order properties

- Using the above simplifications, we first consider the HF gradient

$$E_{0;pq}^{[1]} = \left[\frac{\partial E^{HF}}{\partial \kappa_{pq}} \right]_{\epsilon=0} = \left[\frac{\partial}{\partial \kappa_{pq}} \langle 0 | [\hat{\kappa}, \hat{H}] | 0 \rangle \right]_{\epsilon=0} = \left[\frac{\partial}{\partial \kappa_{pq}} \langle 0 | \hat{H}^{\{1\}} | 0 \rangle \right]_{\epsilon=0} = \frac{\partial}{\partial \kappa_{pq}} \langle 0 | \hat{H}_0^{\{1\}} | 0 \rangle$$

- Actual expression

$$E_{0;ai}^{[1]} = \left[\frac{\partial E^{HF}}{\partial \kappa_{ai}} \right]_{\epsilon=0} = -F_{ia}$$

- First-order properties:

$$\left[\frac{dE^{HF}}{d\varepsilon_A} \right]_{\epsilon=0} = \left[\frac{\partial E^{HF}}{\partial \varepsilon_A} \right]_{\epsilon=0} = \left[\frac{\partial}{\partial \varepsilon_A} \langle 0 | \hat{H} | 0 \rangle \right]_{\epsilon=0} = \langle 0 | \hat{h}_A | 0 \rangle$$

Hartree-Fock response theory

Second-order molecular properties

- First-order response equation

$$E_0^{[2]} \mathbf{X}_A^{[1]} = -\mathbf{E}_A^{[1]}$$

- Solution vector:

$$\mathbf{X}_A^{[1]} = \begin{bmatrix} X \\ X^* \end{bmatrix}; \quad X_{ai} = \kappa_{ai}^A$$

- Property gradient:

$$\mathbf{E}_A^{[1]} = \begin{bmatrix} g \\ g^* \end{bmatrix}; \quad g_{ai} = -h_{ia}$$

- Electronic Hessian

$$E_0^{[2]} = \begin{bmatrix} A & B \\ B^* & A^* \end{bmatrix}; \quad \begin{aligned} A_{ai,bj} &= \left[\frac{\partial^2 E^{HF}}{\partial \kappa_{ai}^* \partial \kappa_{bj}} \right]_{\epsilon=0} = \delta_{ij} F_{ab} - \delta_{ab} F_{ji} + (ai||jb) \\ B_{ai,bj} &= \left[\frac{\partial^2 E^{HF}}{\partial \kappa_{ai}^* \partial \kappa_{bj}^*} \right]_{\epsilon=0} = (ai||bj) \end{aligned}$$

- Second-order molecular properties

$$\left[\frac{d^2 E^{HF}}{d\epsilon_A d\epsilon_B} \right]_{\epsilon=0} = E_{AB}^{[0]} - \mathbf{E}_A^{[1]\dagger} \left(E_0^{[2]} \right)^{-1} \mathbf{E}_B^{[1]} = \langle 0 | \hat{h}_{AB} | 0 \rangle + \mathbf{E}_A^{[1]} \mathbf{X}_B^{[1]}$$

- Can we somehow extract excitation energies and transition moments ?

▶ Yes, we can !

Hartree-Fock response theory

The frequency-dependent linear response function

- Linear response function in the static case

$$\langle\langle \hat{A}; \hat{B} \rangle\rangle_0 = -\mathbf{E}_A^{[1]\dagger} \left(E_0^{[2]} \right)^{-1} \mathbf{E}_B^{[1]}$$

- Generalization to dynamics properties

$$\langle\langle \hat{A}; \hat{B} \rangle\rangle_\omega = -\mathbf{E}_A^{[1]\dagger} \left(E_0^{[2]} - \omega S_0^{[2]} \right)^{-1} \mathbf{E}_B^{[1]},$$

- where appears the generalized metric

$$S_0^{[2]} = \begin{bmatrix} \Sigma & \Delta \\ -\Delta^* & -\Sigma^* \end{bmatrix}; \quad \begin{aligned} \Sigma_{ai,bj} &= -\delta_{ab}\delta_{ij} \\ \Delta_{ai,bj} &= 0 \end{aligned}$$

Hartree-Fock response theory

Excitation energies and transition moments from the linear response function

- We start from

$$\langle\langle \hat{A}; \hat{B} \rangle\rangle_{\omega} = -\mathbf{E}_A^{[1]\dagger} \left(E_0^{[2]} - \omega S_0^{[2]} \right)^{-1} \mathbf{E}_B^{[1]}$$

- We insert a non-singular matrix \mathbf{X}

$$\begin{aligned} \langle\langle \hat{A}; \hat{B} \rangle\rangle_{\omega} &= -\mathbf{E}_A^{[1]\dagger} \mathbf{X} \mathbf{X}^{-1} \left(E_0^{[2]} - \omega S_0^{[2]} \right)^{-1} \left(\mathbf{X}^{\dagger} \right)^{-1} \mathbf{X}^{\dagger} \mathbf{E}_B^{[1]} \\ &= -\mathbf{E}_A^{[1]\dagger} \mathbf{X} \left(\mathbf{X}^{\dagger} E_0^{[2]} \mathbf{X} - \omega \mathbf{X}^{\dagger} S_0^{[2]} \mathbf{X} \right)^{-1} \mathbf{X}^{\dagger} \mathbf{E}_B^{[1]} \end{aligned}$$

- The resolvent matrix $\left(\mathbf{X}^{\dagger} E_0^{[2]} \mathbf{X} - \omega \mathbf{X}^{\dagger} S_0^{[2]} \mathbf{X} \right)$ can be brought to diagonal form by solving the generalized eigenvalue problem

$$E_0^{[2]} \mathbf{X}_n - \omega_n S_0^{[2]} \mathbf{X}_n = 0,$$

giving approximate excitation energies $\hbar\omega_n$.

- Corresponding approximate transition moments are obtained as $\mathbf{E}_A^{[1]\dagger} \mathbf{X}_n$.
- Note that we obtain these quantities without explicit calculation of excited states !

- From classical mechanics we saw that external fields are introduced into the Hamiltonian through the substitutions

$$H(\mathbf{r}, \mathbf{p}, t) = H_p(\mathbf{r}, \boldsymbol{\pi}, t) + q\phi(\mathbf{r}, t); \quad \mathbf{p} = \boldsymbol{\pi} + q\mathbf{A}$$

- The potentials have gauge freedom

$$\mathbf{A} \quad \rightarrow \quad \mathbf{A}' = \mathbf{A} + \nabla\chi$$

$$\phi \quad \rightarrow \quad \phi' = \phi - \partial_t\chi$$

- Similarly, in the electronic QM Hamiltonian, minimal substitution gives

$$\hat{H}(\hat{\mathbf{r}}, \hat{\mathbf{p}}, t) = \hat{H}_p(\hat{\mathbf{r}}, \hat{\boldsymbol{\pi}}, t) - e\phi(\mathbf{r}, t); \quad \hat{\mathbf{p}} = \hat{\boldsymbol{\pi}} - e\mathbf{A}$$

- Gauge transformations may be induced by **local** unitary transformations

- Suppose that we have a wave function satisfying the time-dependent wave equation

$$\left(\hat{H} - i\hbar\partial_t\right) \psi(\mathbf{r}, t) = 0$$

- We consider a unitary transformation of our equation

$$\hat{U}^{-1} \left(\hat{H} - i\hbar\partial_t\right) \hat{U} \hat{U}^{-1} \psi(\mathbf{r}, t) = \hat{U}^{-1} \left(\hat{H} - i\hbar\partial_t\right) \hat{U} \psi'(\mathbf{r}, t) = 0$$

- We choose a local unitary transformation on the form

$$\hat{U} = \exp \left[-\frac{i}{\hbar} e\chi(\mathbf{r}, t) \right]$$

- We observe the following:

- ▶ $i\hbar\partial_t \hat{U} \psi'(\mathbf{r}, t) = \hat{U} [e\partial_t \chi(\mathbf{r}, t) + i\hbar\partial_t] \psi'(\mathbf{r}, t)$
- ▶ $\hat{\pi} \hat{U} \psi'(\mathbf{r}, t) = (-i\hbar\nabla + e\mathbf{A}) \hat{U} \psi'(\mathbf{r}, t) = \hat{U} [\hat{\pi} - e\nabla\chi(\mathbf{r}, t)] \psi'(\mathbf{r}, t)$

- We conclude that

$$\left(\hat{H} - i\hbar\partial_t\right) \psi(\mathbf{r}, t) = \left(\hat{H}_p(\hat{\mathbf{r}}, \hat{\boldsymbol{\pi}}, t) - e\phi(\mathbf{r}, t) - i\hbar\partial_t\right) \psi(\mathbf{r}, t) = 0$$

- becomes

$$\hat{U}^{-1} \left(\hat{H} - i\hbar\partial_t\right) \hat{U}\psi'(\mathbf{r}, t) = \left(\hat{H}_p(\hat{\mathbf{r}}, \hat{\boldsymbol{\pi}}', t) - e\phi'(\mathbf{r}, t) - i\hbar\partial_t\right) \psi'(\mathbf{r}, t) = 0,$$

- with

$$\begin{aligned} \hat{\mathbf{A}}' &= \mathbf{A} - \nabla\chi \\ \hat{\boldsymbol{\pi}}' &= \hat{\boldsymbol{\pi}} + e\hat{\mathbf{A}}'; \\ \phi' &= \phi + \partial_t\chi \end{aligned}$$

- In principle calculated observable are invariant under the gauge transformation, e.g.

$$\langle \psi' | \hat{\Omega} | \psi' \rangle = \langle \psi | \hat{U} \hat{\Omega} \hat{U}^{-1} | \psi \rangle = \langle \psi | \hat{\Omega} | \psi \rangle,$$

<5->..but this may not be the case in a finite basis.

Looking at a gauge-transformed wave function

(Figure from Trygve Helgaker)

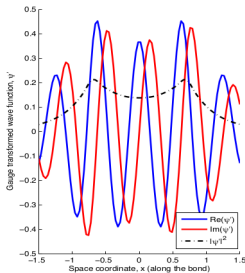
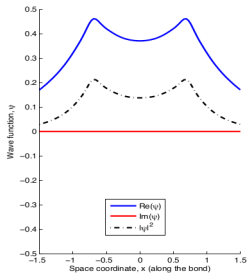
- Consider the vector potential of a uniform magnetic field

$$\mathbf{A}_O(\mathbf{r}_i) = \frac{1}{2}(\mathbf{B} \times \mathbf{r}_{iO})$$

- The gauge origin may be shifted by a gauge transformation

$$\mathbf{A}_G(\mathbf{r}_i) = \frac{1}{2}(\mathbf{B} \times \mathbf{r}_{iG}) = \mathbf{A}_O(\mathbf{r}_i) - \nabla_i \chi(\mathbf{r}_i); \quad \chi(\mathbf{r}_i) = \mathbf{r}_i \cdot \mathbf{A}_O(\mathbf{G})$$

- Illustration:** H_2 on the z -axis with a magnetic field $B = 0.2$ a.u. in the y -direction
 - wave function with a gauge origin at $O = (0, 0, 0)$ (left) and $G = (100, 0, 0)$ (right)



- Is there are optimal gauge origin when calculating magnetic properties ?
 - ▶ For atoms: yes
 - ▶ For molecules: no
- One option is to introduce London orbitals,
 - ▶ also called Gauge-Including Atomic Orbitals (GIAOs)

$$\chi_{\mu}(\mathbf{r}) \rightarrow \omega_{\mu}(\mathbf{r}) = \exp\left[-\frac{i}{\hbar} e\mathbf{r} \cdot \mathbf{A}_G(\mathbf{R}_{\mu})\right] \chi_{\mu}(\mathbf{r})$$

- ▶ removes dependence on some arbitrary gauge origin \mathbf{G} by shifting the gauge origin to the center of the basis function

Dissociation with and without London orbitals

(Figure from Trygve Helgaker)

- Let us consider the **FCI dissociation of H_2** in a magnetic field
 - ▶ **full lines:** with London atomic orbitals
 - ▶ **dashed lines:** without London orbitals

